

# Large Synoptic Survey Telescope

# The LSST CCD Development Program

I. Kotov (BNL), J. Frank (BNL), J. Geary (SAO), D.K. Gilmore (SLAC), P. O'Connor (BNL), V. Radeka (BNL), P. Takacs (BNL), J. A. Tyson (UCD)

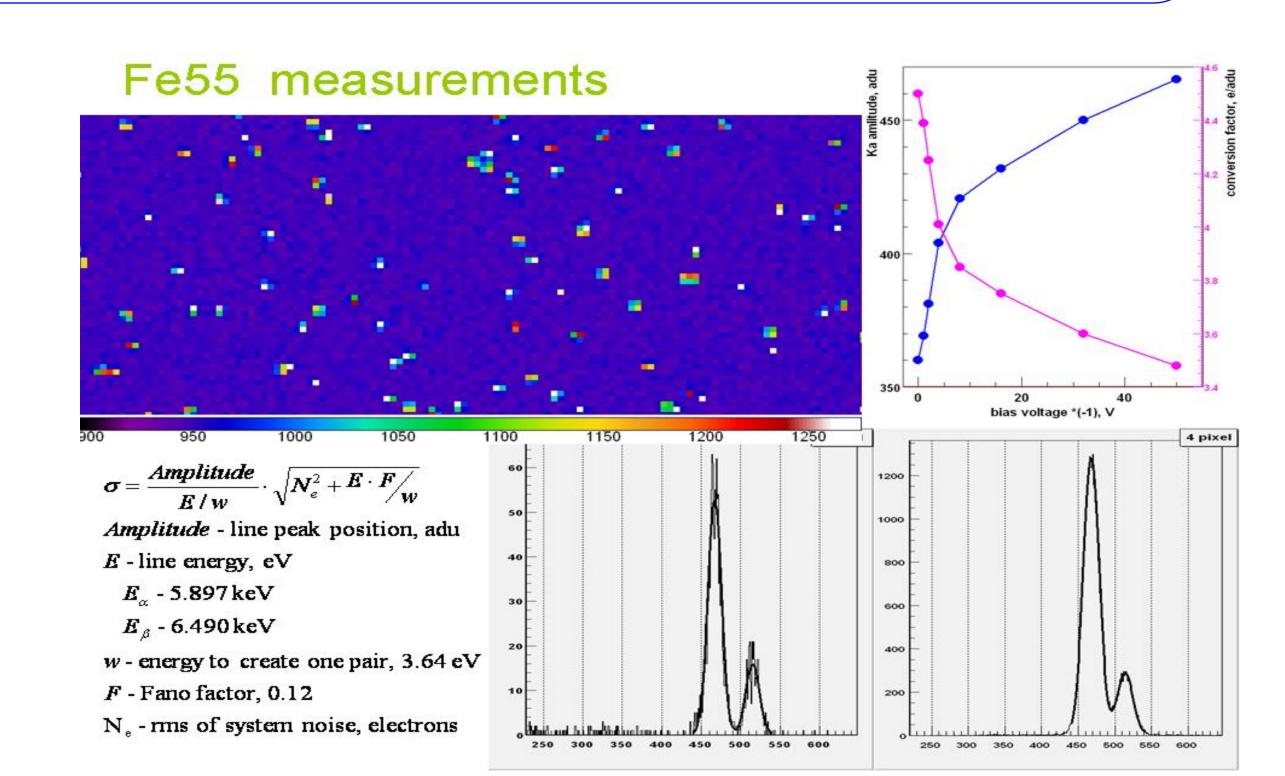
The LSST focal plane array (FPA) will be the largest ever made. The sensors must produce low read noise, high QE in the red, and a very tight PSF. This will all be necessary to do the science at the LSST. The principle underlying the development plan is that for an FPA involving about 200 large format (4k x 4k) sensors, an industrial approach has to be developed and adopted. In this initial phase of CCD development, we have targeted specific technology challenges at competitively selected vendors, with the goal of establishing both the technical characteristics of actual sensors, based on our projected requirements, and the industrial feasibility of their production. The CCD technology challenges we have targeted in particular are over-depleted high resistivity devices in the 100 micron thickness range with a biased conductive window. Initial test results from the first devices in a smaller format resulting from this study program will be presented, demonstrating that these challenges can be overcome...

### CCD sensors development program

- industrial approach
  - study devices from industrial vendors are already available for testing
- technology challenges
  - thick ~100um devices with biased conductive window
  - high resistivity silicon substrate
  - over-depletion to create high E field
- sensor characterization
  - optimized setups for specific tests
  - data acquisition automation
  - data analysis packages

#### Sensor characterization procedures

- QE measurements
  - monochromator setup
- PSF measurement techniques development:
  - virtual knife edge technique
  - modulation transfer function MTF
- Fe55 measurements
  - gain, CTE, noise
- Dark current and defects characterization
- Linearity and Well capacity
- Afterglow (residual image)
- Flatness measurement station



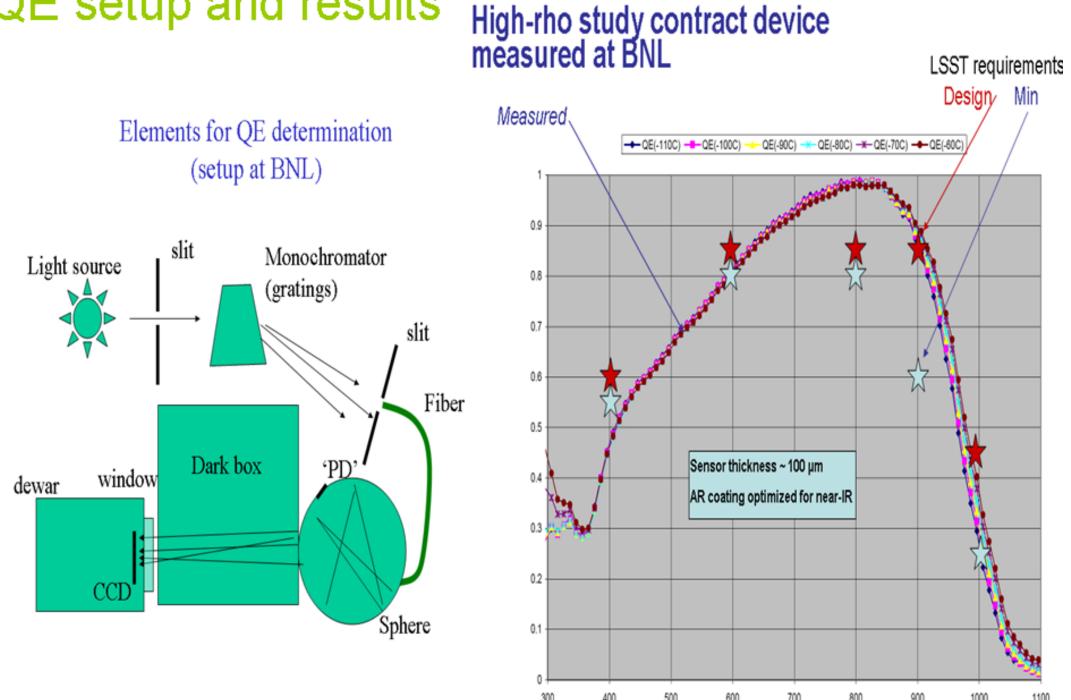
### Sensors key performance requirements:

- Quantum Efficiency (QE)
- Sensor Quantum Efficiency Wavelength | Design | Min | Unit Point Spread Function (PSF) 85 60 % 45 25 % PSF << 0.7" (0.2") 1000 nm
  - $\rightarrow$  pixel size (0.2" = 10  $\mu$ m) → high internal field in the sensor → bias voltage -50V
- → high resistivity substrate (> 5 kOhm·cm) Dark current and defects

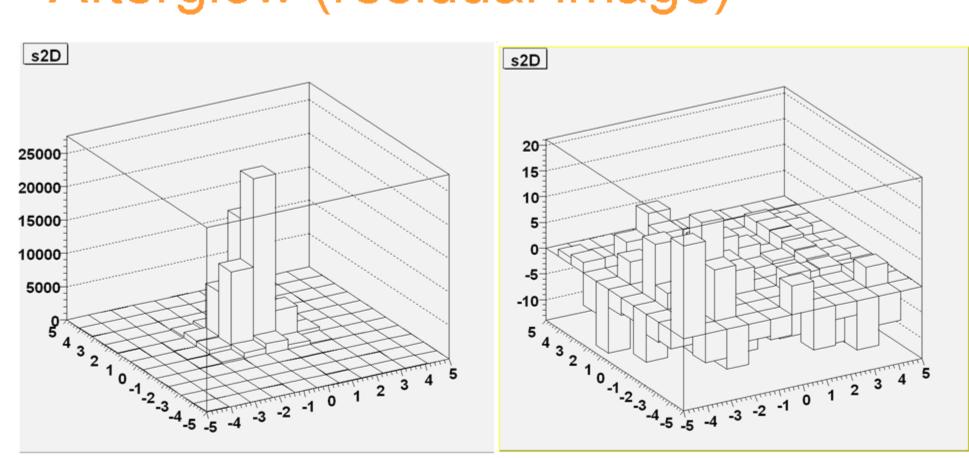
etc.....

 Afterglow (residual image) Gain, Linearity, Noise, Crosstalk, Charge Transfer Efficiency CTE, Flatness, Operating temperature, Temperature uniformity & stability,

# QE setup and results

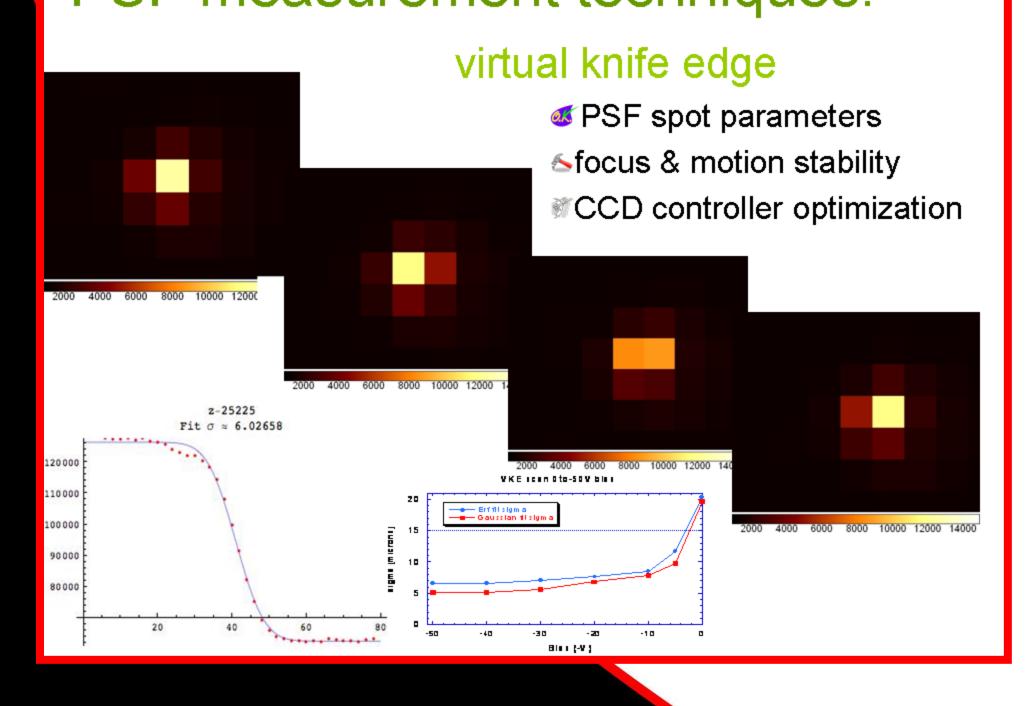


# Afterglow (residual image)

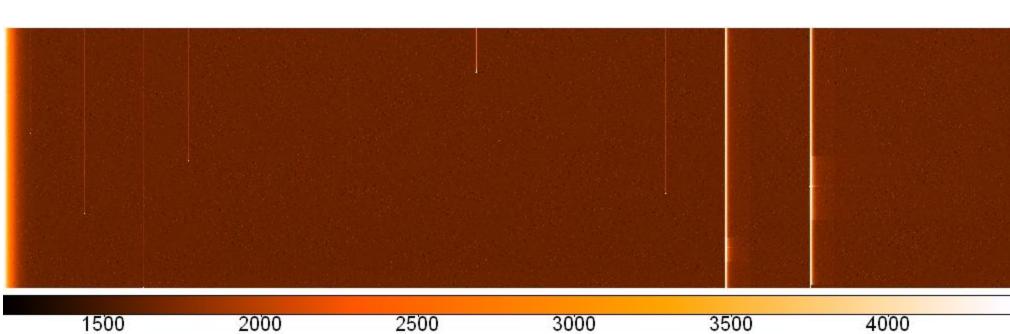


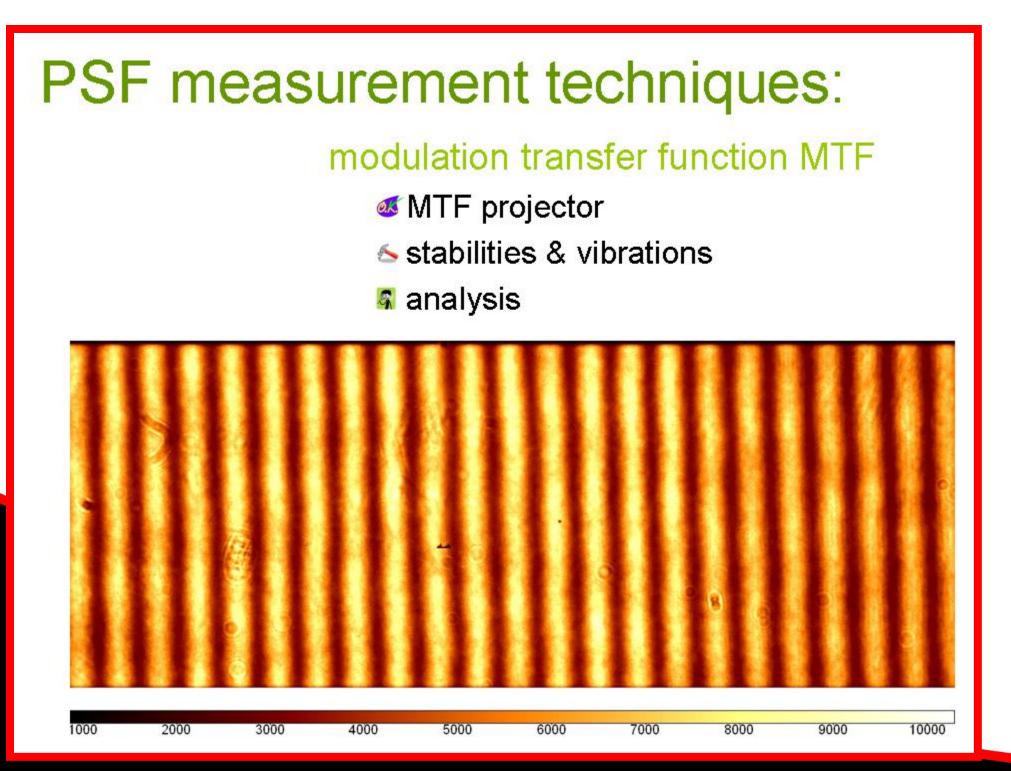
- bias subtracted using zero exposures
- no "after glow" signal on the level of 0.03% (measurement accuracy)

# PSF measurement techniques:

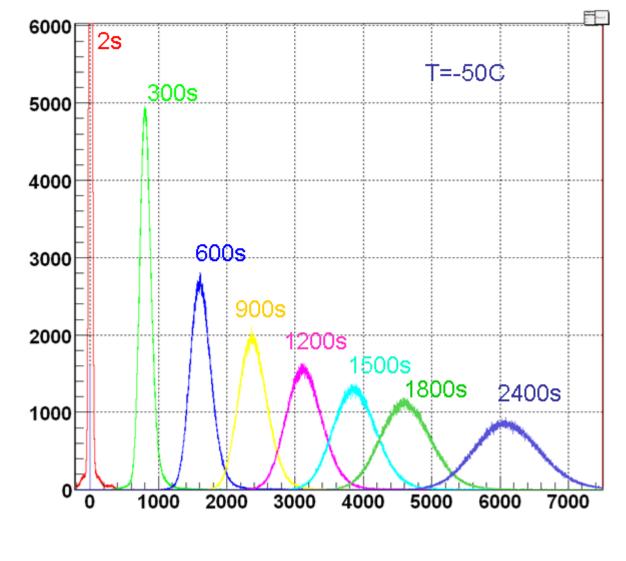


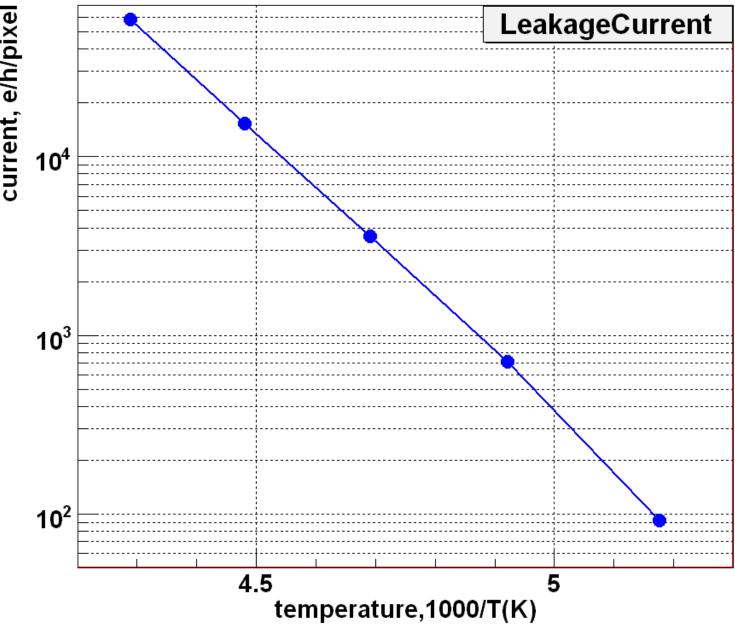
# Leakage Current measurements





# Leakage Current





LSST is a public-private partnership. Design and development activity is supported by in part the National Science Foundation under Scientific Program Order No. 9 (AST-0551161) and Scientific Program Order No. 1 (AST-0244680) through Cooperative Agreement AST-0132798. Portions of this work are supported by the Department of Energy under contract DE-AC02-76SF00515 with the Stanford Linear Accelerator Center, contract DE-AC02-98CH10886 with Brookhaven National Laboratory, and contract W-7405-ENG-48 with Lawrence Livermore National Laboratory. Additional funding comes from private donations, grants to universities, and in-kind support at Department of Energy laboratories and other LSSTC Institutional Members.